

### **CASSINI T122 SEGMENT**

**Rev 239 Handoff Package** 

Segment Boundary 2016-222T22:10:00 - 2016-224T12:25:00

31 DEC2015

Karl Mitchell

SMT report and SPASS
Science Highlights
Notes & Liens

This document has been reviewed and determined not to contain export controlled technical data

## **SMT** report

DATA VOLUME SUMM	RY TRANSFEI	R FRAME OVERHEA	D INCLUDED	(80 BITS	PER	8800-BIT FRAME)
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			l		OBS	ERVATIO	ON_PERIO	DD					DOWNLINK_PASS			
						P4			P5	RECO	ORDED		PLAYE	3ACK		
DOWNLINK PASS NAME	Start doy hh:mm	End doy hh:mm	START (Mb)		HK+E (Mb)	TOTAL (Mb)	CPACTY (Mb)	MRGN (Mb)	OPNAV (Mb)	SCI (Mb)	ENGR (Mb)	TOTAL (Mb)	CPACTY MARGN (Mb) (Mb)	_	ARGN (%)	CAROVR (Mb)
SP_239EA_C34BWGNON223_PRIME SP_239EA_M34BWGNON223_PRIME	223 06:10	223 14:24 223 20:30	1400	13 0 0 668	4 0 0 29		3322 3322 3322 3322	1795 1922 585 10	0 0 0 0	92 1599 73 199	41 49 36 53	1660 3048 2846 3564	260 -1401 311 -2737 231 -2615 3764 199	10 10 10 200	0% 0% 0% 5%	2737 2615

SSR PARTITION SIZE SUMMARY - SELECTED SSR CONFIGURATION: DOUBLE

		SSR A/B		
OBSERVATION PERIOD	P4 Size (Frames)	P5 Size (Frames)	P6 Size (Frames)	
	* I 188954	NITIAL CONDITI 10	ON * 38863	
SP_239NA_OBSERV227_NA SP_239NA_OBSERV223_NA	188954 188954	10 10	38863 38863	

DATA VOLUME REPORT --- TRANSFER FRAME OVERHEAD NOT INCLUDED

Event	Start doy hh		nd oy	hh:mm	CAPS (Mb)	CDA (Mb)	CIRS (Mb)	INMS (Mb)	ISS (Mb)	MAG (Mb)	MIMI (Mb)	RADAR (Mb)	RPWS (Mb)	UVIS (Mb)	VIMS (Mb)	PROBE (Mb)	ENGR (Mb)	TOTAL (Mb)
OBSERVATION_NOR SP_239EA_G34BWGNON222_PRIME SP_239EA_C34BWGNON223_PRIME SP_239EA_M34BWGNON223_PRIME DAILY TOTAL SCIENCE	223 06	:13 2 :10 2 :24 2	23 23 23	23:13 06:10 14:24 20:30 20:30	0.0 0.0 0.0 0.0	2.0 13.1 45.7 11.5 72.3	0.0 0.0 0.0 0.0	0.4 2.5 13.0 2.2 18.1	0.0 0.0 0.0 0.0	1.9 21.2 48.5 10.9 82.4	3.2 21.2 34.6 18.7 77.7	0.0	5.0 32.7 1442.7 28.8 1509.3	0.1 0.0 0.0 0.0 0.1	0.0 0.0 0.0 0.0	0.0 0.0 0.0 0.0	4.4 0.0 0.0 0.0 4.4	17.1 90.7 1584.5 72.2
OBSERVATION_NOR SP_239EA_C70METNON224_PRIME DAILY TOTAL SCIENCE	223 20 224 03 223 20	:25 2	24	03:25 12:25 12:25	0.0 0.0 0.0	13.0 17.0 30.0	77.6 86.4 164.0	2.5 3.2 5.7	350.0 0.0 350.0	12.3 16.0 28.3	21.1 27.5 48.7	0.0 0.0 0.0	32.5 42.4 75.0	103.0 4.9 107.9	50.0 0.0 50.0	0.0 0.0 0.0	28.8 0.0 28.8	690.9 197.5

	CAPS (Mb)	CDA (Mb)	CIRS (Mb)	INMS (Mb)	ISS (Mb)	MAG (Mb)	MIMI (Mb)	RADAR (Mb)	RPWS (Mb)	UVIS (Mb)	VIMS (Mb)	PROBE (Mb)
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TOTAL RECORDED (OPNAV data not included)

0.0 102.3

239TI T122   1698 km	20311_1122 1030 KIII
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Start Time	End Time	Prime Activity	Obs. Detail	Op Mode	TLM Mode	Comments
2016-222T22:10:00	2016-222T22:25:00	Observation gap	No activity	RSS_K_RWAF	S_N_ER_3	Recent update to segment boundaries introduced a 15-minute gap
2016-222T22:25:00	2016-222T23:05:00	SP Turn to WP	X-band to Earth/RA-Dec:123/-57	RSS_K_RWAF	S_N_ER_3	Recent change of fly-by start time to 22:10:00 means there is now a 15-minute gap between start and SP Turn to WP.
2016-222T23:05:00	C/A-09:17:15	OD Uncertainty Dead Time		RSS_K_RWAF	S_N_ER_3	
C/A-09:17:15	0	SP/RSS Gravity	Warm up RSS during downlink (TN1b)	RSS_K_RWAF	RTE_N_SPB	CDA Collaborative Rider
2016-223T08:30:53		CLOSEST APPROACH	XBAND to EARTH, NEG_X to NSP (Tc2a)			
0	+12:00:00	SP/RSS Gravity	(TN1b)	RSS_K_RWAF	RTE_N_SPB	CDA Collaborative Rider
+12:00:00	C/A+17:59:07	ISS	(TC1a/TN1a)	DFPW Normal	S_N_ER_3	
C/A+17:59:07	2016-224T02:45:00	OD Uncertainty Dead Time		DFPW Normal	S_N_ER_3	
2016-224T02:45:00	2016-224T03:25:00	SP Turn to Earth for downlink	X-band to Earth/RA-Dec:(0.0,0.0,-9.5 deg. offset)	DFPW Normal	S_N_ER_3	RSS requested change to RSS_K_RWAF from 03:19:50 until 03:25:00.
2016-224T03:25:00	2016-224T12:25:00	Canberra 70M	5-hr rolling then MIMI.NEG_Y to Saturn (0,0,-9.5)	RSS_K_RWAF	RTE_N_SPB	RSS GSE on C34 concurrent. KTPY bias, if needed, at very end of downlink (last 1:30 hr, non-rolling). Switching back to DFPW Normal from 12:23:12 to 12:25:00

# **SPASS**

	SPASS fo	or Delivery: TI239_1	T122 Records 1-12 (Page 1 of 1	1.)		Obs	servation Attitude		
Request	♦ Riders ♦	Start (SCET)	♦ Start (Epoch)	Duration 💠	End (SCET)	Primary	Secondary	<b>♦</b> Comments	<b>\$</b>
Sequence S95, length = 74 days		2016-178T15:44:00		073T18:52:00	2016-252T10:36:00				
Titan Flyby T122 Segment		2016-222T22:10:00		001T14:15:00	2016-224T12:25:00				
SP 239EA WAYPTTURN222 PRIME	<u>R</u>	2016-222T22:25:00		000T00:40:00	2016-222T23:05:00	XBAND to Earth	NEG_X to 123.0/-57.0		
NEW WAYPOINT		2016-222T23:05:00		001T04:20:00	2016-224T03:25:00	XBAND to Earth	NEG_X to 123.0/-57.0		
SP 239EA DEADTIME222 PRIME	<u>M, R</u>	2016-222T23:05:00		000T00:08:38	2016-222T23:13:38	XBAND to Earth	NEG_X to 123.0/-57.0	Collaborative Rider(s): CDA	
SP 239EA G34BWGNON222 PRIME	<u>M, R</u>	2016-222T23:13:38	GMB_E239_TITAN_T122-000T09:17:15	000T06:56:22	2016-223T06:10:00	XBAND to Earth	NEG_X to 123.0/-57.0	Collaborative Rider(s): CDA, RSS	
SP 239EA C34BWGNON223 PRIME	<u>M, R</u>	2016-223T06:10:00	GMB_E239_TITAN_T122-000T02:20:53	000T08:14:00	2016-223T14:24:00	XBAND to Earth	NEG_X to 123.0/-57.0	Collaborative Rider(s): CDA, RSS	
239TI (t) T122 TITAN Outbou		2016-223T08:30:53		000T00:00:01	2016-223T08:30:54				
SP 239EA M34BWGNON223 PRIME	<u>M, R</u>	2016-223T14:24:00	GMB_E239_TITAN_T122+000T05:53:07	000T06:06:53	2016-223T20:30:53	XBAND to Earth	NEG_X to 123.0/-57.0	Collaborative Rider(s): CDA, RSS	
ISS 239TI MONITORNA001 PRIME	<u>C, M, U, V</u>	2016-223T20:30:53	GMB_E239_TITAN_T122+000T12:00:00	000T05:59:07	2016-224T02:30:00	ISS_NAC to Titan	NEG_X to 123.0/-57.0	Collaborative Rider(s): CDA. No Preference to secondary pointing	
SP 239EA DEADTIME224 PRIME		2016-224T02:30:00	GMB_E239_TITAN_T122+000T17:59:07	000T00:15:00	2016-224T02:45:00	XBAND to Earth	NEG_X to 123.0/-57.0		
SP 239EA DLTURN224 PRIME		2016-224T02:45:00		000T00:40:00	2016-224T03:25:00	XBAND to Earth (0.0,0.0 deg. offset)	0,-9.5 NEG_Y to Saturn		
NEW WAYPOINT		2016-224T03:25:00		000T09:00:00	2016-224T12:25:00	XBAND to Earth (0.0,0.0 deg. offset)	0,-9.5 NEG_Y to Saturn		
SP 239EA C70METNON224 PRIME	<u>C, E, R</u>	2016-224T03:25:00		000T09:00:00	2016-224T12:25:00	XBAND to Earth (0.0,0.0 deg. offset)	0,-9.5 5_Hr_Rolling	MIMI.NEG_Y to Saturn (0,0,-9.5).SRU.CIRS heating.	



#### Sequence S95/T122: Summary of PIEs and Other High Priority Observations

					Comments (e.g., pointing tolerance,	Science Traceability Matrix	
Discipline	CIMS Request Name	Start Time	End Time	pointing	uniqueness; relative priority)	Code(s)	Pointing designer POC
					RSS Gravity Science; CDA dust		
					occultation experiment (CDA-		
Titan	SP_239EA_G34BWGNON222_F	2016-222T23:13:38	2016-223T06:10:00	Significant Science Impact	friendly secondary).	TN1b	Karl.L.Mitchell@jpl.nasa.gov
					RSS Gravity Science; CDA dust		
					occultation experiment (CDA-		
Titan	SP_239EA_C34BWGNON223_F	2016-223T06:10:00	2016-223T14:24:00	Significant Science Impact	friendly secondary).	TN1b	Karl.L.Mitchell@jpl.nasa.gov
					RSS Gravity Science; CDA dust		
					occultation experiment (CDA-		
Titan	SP_239EA_M34BWGNON223_F	2016-223T14:24:00	2016-223T20:30:53	Significant Science Impact	friendly secondary).	TN1b	Karl.L.Mitchell@jpl.nasa.gov

Collaborative riders: CDA\_239TI\_GRAVITY001\_RSS

DOY 222: RSS will start its last Gravity flyby on HGA, one of a few such critical observations gathered over the course of the mission that contribute data critical for: (1) assessing the presence of a global subsurface ocean by measuring the short-period changes of the gravity field induced by Saturn's tidal field (eccentricity tides); (2) determining the geoid and the presence of large scale gravity anomalies; (3) determine the rheology of the icy crust by correlative analysis with altimetric data. CDA, a collaborative rider, will be using a new "Dust Occultation" technique to observe the interaction of nanodust streams with Titan's atmosphere. Dust particles are used like photons: An obstacle hinders the dust particles from a (more or less) collimated source to reach the observer. T122 is the only and last opportunity to observe a new occultation phenomena first observed during T99 (2014), in which CDA observed a drop out of the impact rate of dust impacts when Saturn vanished behind Titan's atmosphere during the Cassini flyby. The new measurements will constrain the grain parameters like speed, mass and direction they and will provide essential input for the modeling of the interaction of fast nanograin ablation in thin atmospheres.

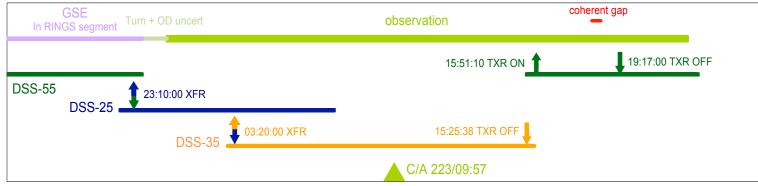
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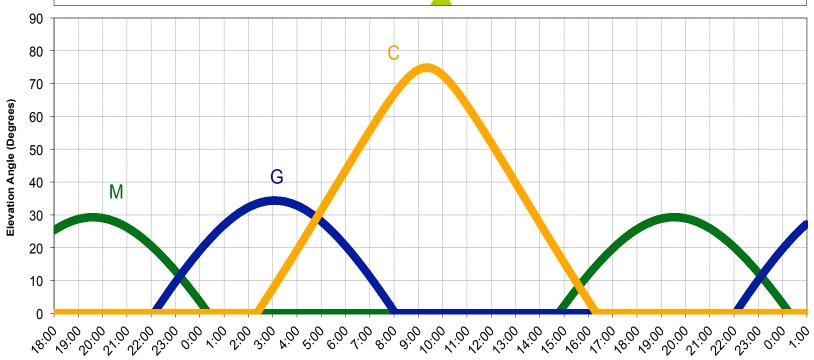
DOY 223: RSS will continue its last Gravity flyby on HGA, throughout the inbound pass and 12 hours into the outbound, one of a few such critical observations gathered over the course of the mission that contribute data critical for: (1) assessing the presence of a global subsurface ocean by measuring the short-period changes of the gravity field induced by Saturn's tidal field (eccentricity tides); (2) determining the geoid and the presence of large scale gravity anomalies; (3) determine the rheology of the icy crust by correlative analysis with altimetric data. CDA, a collaborative rider, will be using a new "Dust Occultation" technique to observe the interaction of nanodust streams with Titan's atmosphere. Dust particles are used like photons: An obstacle hinders the dust particles from a (more or less) collimated source to reach the observer. T122 is the only and last opportunity to observe a new occultation phenomena first observed during T99 (2014), in which CDA observed a drop out of the impact rate of dust impacts when Saturn vanished behind Titan's atmosphere during the Cassini flyby. The new measurements will constrain the grain parameters like speed, mass and direction they and will provide essential input for the modeling of the interaction of fast nanograin ablation in thin atmospheres.

DOY 223 (cont.) From 12 hours outbound, ISS will acquire global-scale imaging of Titan's trailing hemisphere at mid-northern latitudes, providing important constraints for the cloud monitoring campaign. VIMS, CIRS and ISS will ridealong with ISS. VIMS will monitor monitoring cloud activity and looking for specular reflections on lakes near the north polar. UVIS will obtain spectra of Titan's mid atmosphere and stratosphere, searching in particular for nitrogen emission features in the mid atmosphere and stratosphere. CIRS will obtain data to monitor Titan's changing seasons, and giving insights into the stratospheric circulation. And make detailed scans of Titan's atmospheric limb near 50N revealing the vertical structure of temperature and trace gas abundances, such as hydrocarbons and nitriles. These valuable data will be used for comparison with observations of equivalent southern latitudes, which are currently experiencing late Fall (south) instead of spring (north). They will also be compared to views of the north earlier in the mission.

DOY 224: **ISS** will continue to acquire global-scale imaging of Titan's trailing hemisphere at mid-northern latitudes, providing important constraints for the cloud monitoring campaign. VIMS, CIRS and ISS will ridealong with ISS. VIMS will monitor monitoring cloud activity and looking for specular reflections on lakes near the north polar. UVIS will obtain spectra of Titan's mid atmosphere and stratosphere, searching in particular for nitrogen emission features in the mid atmosphere and stratosphere. CIRS will obtain data to monitor Titan's changing seasons, and giving insights into the stratospheric circulation. And make detailed scans of Titan's atmospheric limb near 50N revealing the vertical structure of temperature and trace gas abundances, such as hydrocarbons and nitriles. These valuable data will be used for comparison with observations of equivalent southern latitudes, which are currently experiencing late Fall (south) instead of spring (north). They will also be compared to views of the north earlier in the mission.







Time (ERT-UTC)

RTLT ~2:40

An engineering Y bias window overlaps the last 90 minutes of the DOY 124 Canberra 70m downlink, during the MIMI-fixed secondary period following the 5-hr rolling downlink potion. Due to RSS science (see next slide) TOST is requesting that AACS try to avoid placing a bias in this segment. **SCO has agreed to perform an on-Earth bias if a bias must be placed AND if data volume is tight.** Consequently we are not specifying a plan to cut data volume overflow in the event that this Y bias results in the shortening of the downlink. If SSR overflow occurs, data will be cut from RPWS.

RSS requests no thruster during from the beginning of the inbound GSE 222/21:08:38 (in Rings segment) until the end of the outbound GSE 224/12:25:00, with the critical period being from 222/23:13:38 to 223/20:30:53. If biases are needed during the GSEs, try to place them as early as possible during the inbound GSE or as late as possible during the outbound GSE.

## **Notes**

- Pointing:
  - YGAPs avoided in this segment until near end of final downlink due to RSS science.
- Data Volume:
  - Begins with 1510Mb from previous segment (RINGS 239smt 151208.init).
  - SCO will use on-Earth bias if data volume is tight.
  - Constant low downlink rate used during RSS activities to simplify implementation in the event of DSN changes.
  - SIP leads could implement a telemetry wedding cake if more data volume needed
- DSN:
  - Level 3 requests from 2016-222/1600 to 2016-223/2230. Stations: DSS-55, DSS-25, DSS-35.
- Resource checker:
  - Gap in first 15 minutes of segment (not reported in Resource Checker) due to late change to segment boundary. No adverse consequences.
  - ENGR bias during rolling downlink takes place after 5-hr rolling downlink ends: large data buffer non-problematic.
- Opmodes:
  - RSS warmup begins in previous segment.
- Hydrazine:
  - Flyby on RWA.
- **Special Activities:**

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RSS Gravity flyby (SPLAT item)



### Sequence Liens (should all be SPLAT items):

Request for no biases during RSS gravity experiments (SPLAT item)